

# Ray Optics



**CBSE CLASS NOTES**

**SIMIL RAHMAN**

**MSc, D. Electrical Electronics  
Engineering, MA. Eng  
, B. Ed, SET, PhD (Doing)**

**M.E.S Doha, Qatar**

# Ray Optics and Optical Instruments.

## Optics

It is the branch of physics which deals with the study of Nature, production and propagation of light.

### Branches

#### (1) Ray or geometrical optics

It concerns itself with the particle nature of light and is based on

- (i) Rectilinear propagation of light
- (ii) the laws of reflection and refraction of light.

#### (2) wave or Physical Optics

It concerns itself with the wave nature of light and is based on the phenomena like

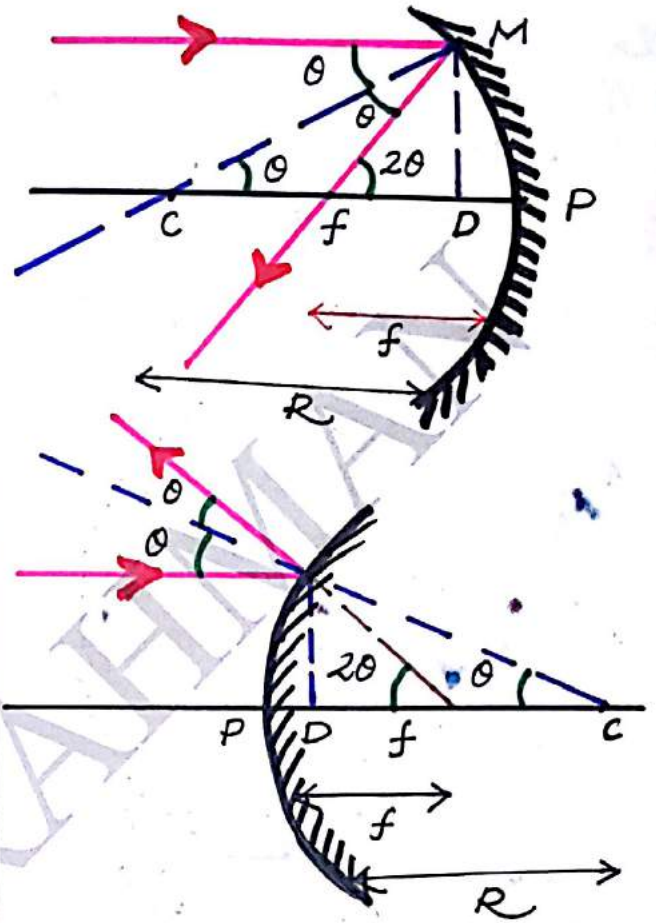
- (i) Interference
- (ii) diffraction
- (iii) Polarisation

### laws of reflection

- (i)  $\angle i = \angle r$
- (ii) Incident ray, reflected ray and normal all lie in the same plane.



## Relation b/w $f$ and $R$



$$\tan \theta = \frac{MD}{CD}$$

$$\tan 2\theta = \frac{MD}{DF}$$

for small angles  
 $\tan \theta = \theta$   
 $\tan 2\theta = 2\theta$

$$\therefore \theta = \frac{MD}{CD} \dots \text{--- (1)}$$

$$2\theta = \frac{MD}{DF} \dots \text{--- (2)}$$

sub (1) in (2)

$$2 \frac{MD}{CD} = \frac{MD}{DF}$$

$$\frac{2}{CD} = \frac{1}{DF}$$

$$\therefore \frac{2}{R} = \frac{1}{f}$$

$$\begin{aligned} CD &\approx CP \\ CP &= R \\ \therefore CD &= R \\ DF &= f \end{aligned}$$

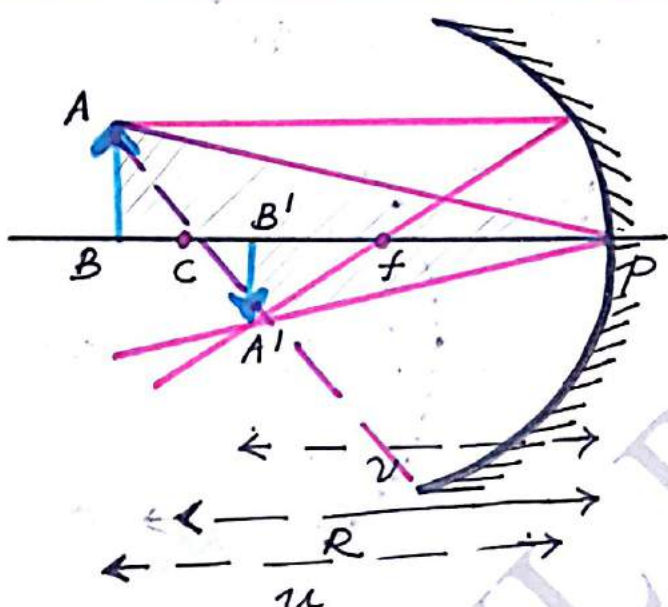
$$R = 2f$$

# Spherical Mirrors

A spherical mirror is a reflecting surface which forms part of hollow sphere.

convex  
concave

Mirror formula -  
concave mirror real image



$\Delta ABC$  and  $\Delta A'B'C$  are similar

$$\frac{AB}{A'B'} = \frac{CB}{CB'} \dots \textcircled{1}$$

$\Delta ABP$  and  $\Delta A'B'P$  similar

$$\frac{AB}{A'B'} = \frac{PB}{PB'} \dots \textcircled{2}$$

$$\textcircled{1} = \textcircled{2}$$

$$\frac{CB}{CB'} = \frac{PB}{PB'}$$

$$\frac{PB - PC}{PC - PB'} = \frac{PB}{PB'}$$

$$\frac{-u - (-R)}{-R - (-v)} = \frac{-u}{-v}$$

$$\frac{-u + R}{-R + v} = \frac{u}{v}$$

$$-uV + VR = -UR + uV \dots \textcircled{3}$$

$$\textcircled{3} \div uVR$$

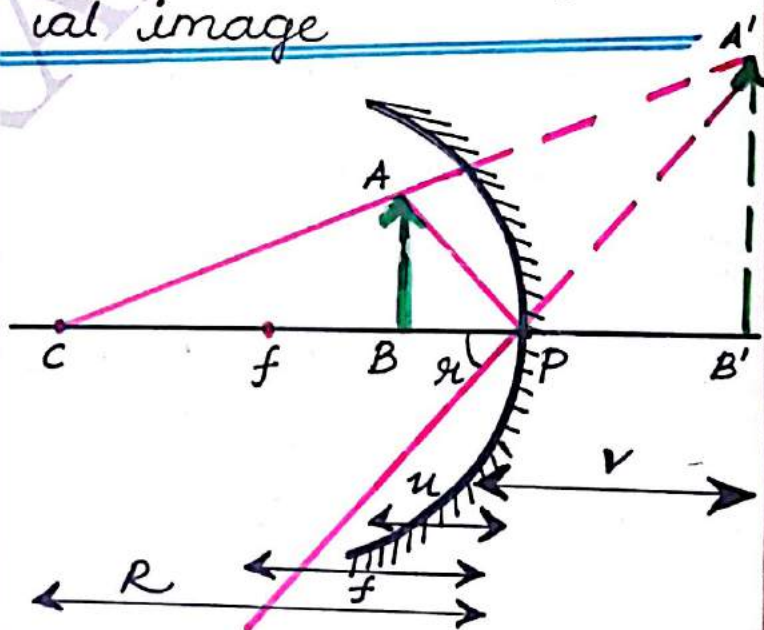
$$-\frac{1}{R} + \frac{1}{u} = -\frac{1}{v} + \frac{1}{R}$$

$$\frac{1}{v} + \frac{1}{u} = \frac{2}{R}$$

but  $\frac{2}{R} = \frac{1}{f}$

$$\frac{1}{v} + \frac{1}{u} = \frac{1}{f} \dots \text{mirror formula}$$

Mirror formula -  
concave mirror - virtual image



$\Delta ABC$  and  $\Delta A'B'C$  similar

$$\frac{AB}{A'B'} = \frac{CB}{CB'} \dots \textcircled{1}$$

$\Delta ABP$  and  $\Delta A'B'P$  similar

$$\frac{AB}{A'B'} = \frac{PB}{PB'} \dots \textcircled{2}$$

$$\textcircled{1} = \textcircled{2}$$

$$\frac{CB}{CB'} = \frac{PB}{PB'}$$

$$\frac{PC - PB}{PC + PB'} = \frac{PB}{PB'}$$

$$\frac{-R - (-u)}{-R + v} = \frac{-u}{v}$$

$$\frac{-R + u}{-R + v} = \frac{-u}{v}$$

$$-Rv + uv = uR - uv \dots (3)$$

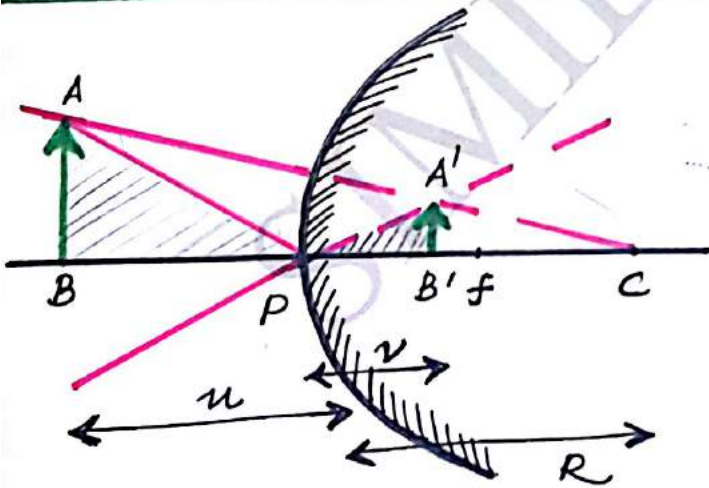
$$(3) \div uvr$$

$$-\frac{1}{u} + \frac{1}{R} = \frac{1}{v} - \frac{1}{R}$$

$$\frac{1}{v} + \frac{1}{u} = \frac{2}{R} \quad \frac{2}{R} = \frac{1}{f}$$

$$\frac{1}{v} + \frac{1}{u} = \frac{1}{f}$$

Mirror formula -  
convex mirror - virtual  
image



$\Delta ABC$  and  $\Delta A'B'C$  - similar

$$\frac{AB}{A'B'} = \frac{CB}{CB'} \dots (1)$$

$\Delta ABP$  &  $\Delta A'B'P$  - similar

$$\frac{AB}{A'B'} = \frac{PB}{PB'} \dots (2)$$

$$(1) = (2)$$

$$\frac{CB}{CB'} = \frac{PB}{PB'} \quad (2)$$

$$\frac{PB + PC}{PC - PB'} = \frac{PB}{PB'}$$

$$\frac{-u + R}{R - v} = \frac{-u}{v}$$

$$-uv + vR = -uR + uv \dots (3)$$

$$(3) \div uvr$$

$$-\frac{1}{R} + \frac{1}{u} = \frac{-1}{v} + \frac{1}{R}$$

$$\frac{1}{v} + \frac{1}{u} = \frac{2}{R} \quad \frac{2}{R} = \frac{1}{f}$$

$$\frac{1}{v} + \frac{1}{u} = \frac{1}{f} \rightarrow \text{mirror formula}$$

Magnification  
- Mirror

$$m = \frac{h_i}{h_o} = \frac{-v}{u} = \frac{f}{f-u} = \frac{f-v}{f}$$

Applications of convex  
and concave mirrors

concave mirror:- Dentists,  
torch light, shaving  
mirror.

convex mirror:- Rear  
view mirror, reflection  
in street lamps.

REFRACTION

The phenomenon of the change in

the path of light as it passes obliquely from one transparent medium to another is called refraction of light.

### Laws of refraction

1, Incident ray, refracted ray, normal all lie in the same plane

2, Snell's law

$$\frac{\sin i}{\sin r} = n_{21} = \frac{n_2}{n_1}$$

### Absolute Refractive Index

$$\mu = \frac{c}{v} = \frac{v_{\text{air}}}{v_{\text{med}}} = \frac{\lambda_{\text{air}}}{\lambda_{\text{med}}}$$

$\mu$  = velocity of light in vacuum

velocity of light in medium

$$\therefore \mu = \frac{c}{v} = \frac{\lambda_{\text{air}}}{\lambda_{\text{medium}}}$$

write expressions for the refractive index of second medium w.r.t refractive index of first medium in terms of speed,  $\lambda$  and refractive index

$$n_{21} = \frac{n_2}{n_1} = \frac{v_1}{v_2} = \frac{\lambda_1}{\lambda_2} = \frac{1}{n_{12}}$$

constant quantities in refraction

(a) frequency

(b) Energy  $E = h\nu$

write physical quantities which changes when a light ray passes from one medium to another medium

$\lambda$  and speed will change.

factors on which refractive index depend upon

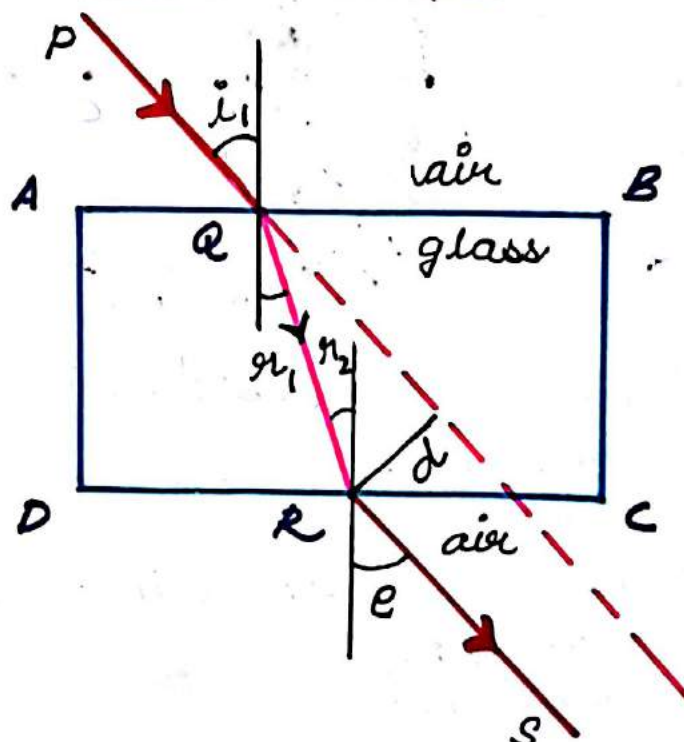
(a) nature of medium

(b)  $\lambda$  of light

when a light ray passes through parallel sides of a glass slab

(i) show that  $\angle i = \angle e$

(ii) Give expression for lateral shift.



Let ABCD be a glass prism, PQ - Incident ray  
 QR - refracted ray  
 RS - emergent ray.

Due to refraction at face AB

$$\mu = \frac{\sin i_1}{\sin r_1} \dots \dots \textcircled{1}$$

Due to refraction at face CD

$$\mu = \frac{\sin e}{\sin r_2} \dots \dots \textcircled{2}$$

from ① and ②

$$i_1 = e$$

$$r_1 = r_2$$

$$\therefore \angle i = \angle e$$

$$\angle r_1 = \angle r_2$$

\* Emergent ray is parallel to incident ray, but it is laterally displaced

\* expression for lateral shift

$$d = \frac{t}{\cos r} \sin (i - r)$$

### Lateral Shift

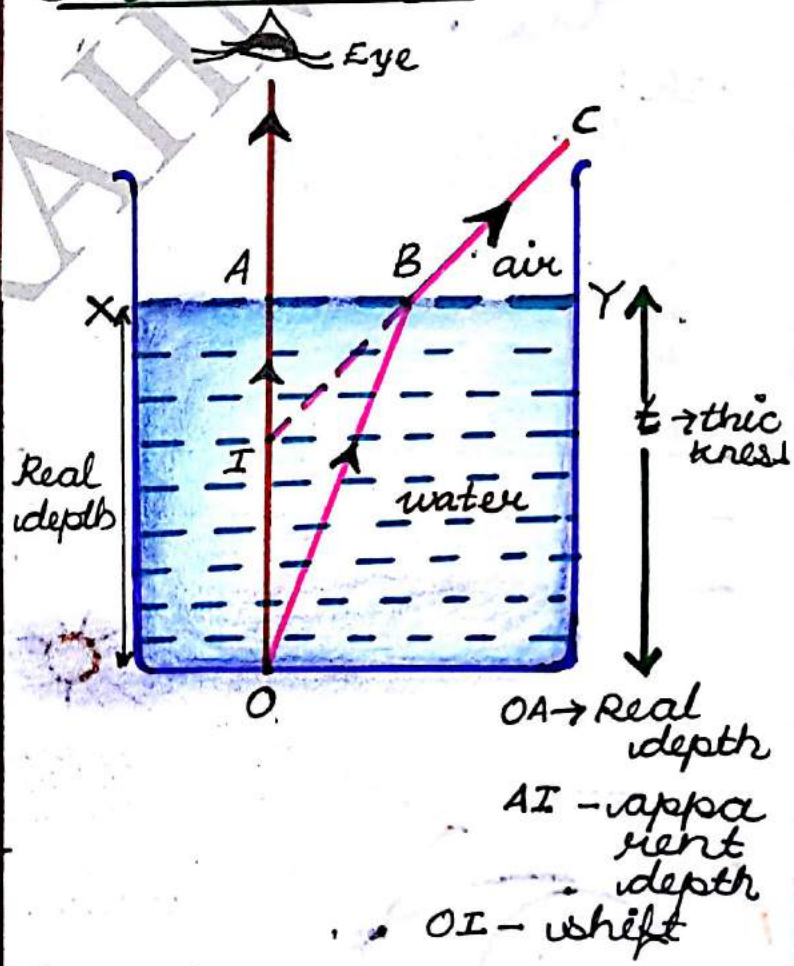
The perpendicular distance between the direction of incident ray and emergent ray, when light ray passes through parallel sides of a glass slab.

### \* lateral shift Increases with

- (1) increase in thickness of glass slab
- (2) Increase of angle of Incidence
- (3) Increase of refractive Index of glass slab.

### Application of Refraction

#### Normal Shift



eg; coins appears to be raised.

$$\mu = \frac{\text{real depth}}{\text{apparent depth}}$$

$$\mu = \frac{OA}{AI}$$

$$AI = \frac{OA}{\mu} \dots \dots \textcircled{1}$$

\* [if the observer is in rarer medium and object is in denser medium]

Shift in position

$$S = OI = OA - AI \dots \dots \textcircled{2}$$

① in ②

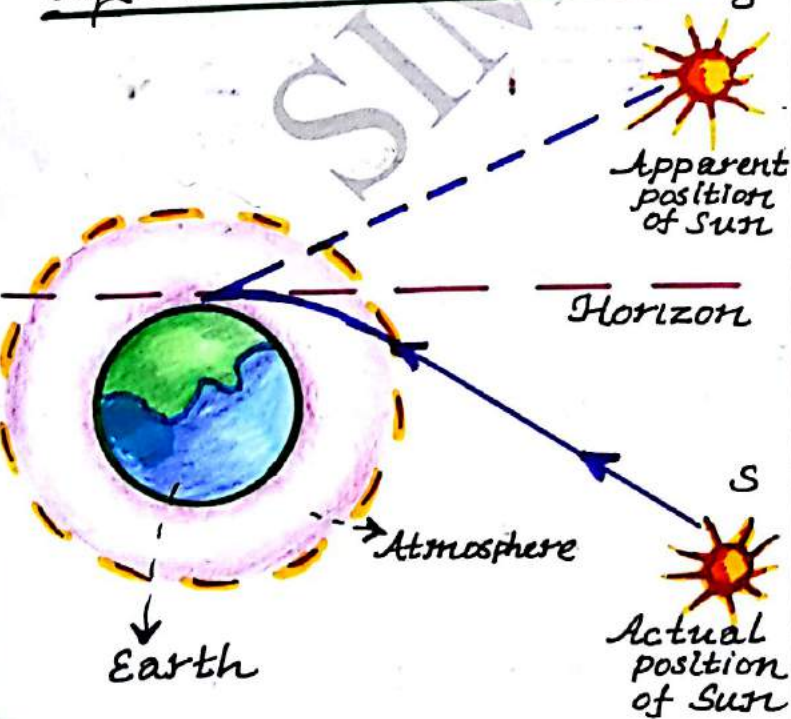
$$S = OA - \frac{OA}{\mu} = OA \left(1 - \frac{1}{\mu}\right)$$

$$S = t \left(1 - \frac{1}{\mu}\right)$$

Note:

$\mu = \frac{\text{app depth}}{\text{Real depth}}$  if observer is in denser medium and object is rarer medium

\* Sun is visible before actual sunrise and after actual sunset.



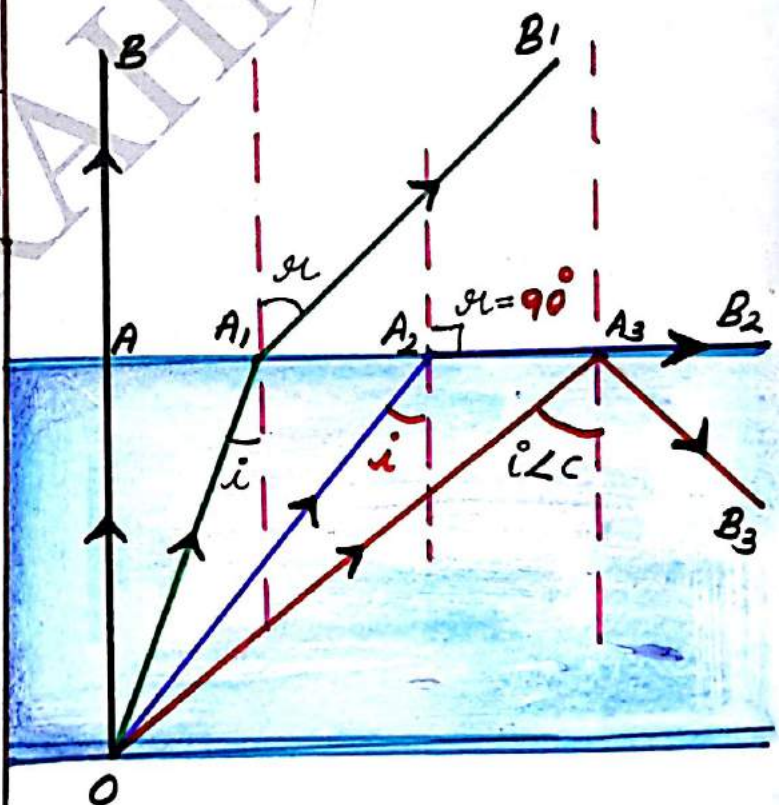
\* Before sunrise and after sunset - sun 's' is below the horizon.

\* Sun light bends towards the normal as it travels from sun (rarer) to earth (denser) to reach observer

\* Due to this atmospheric refraction sun appears above horizontal.

### TIR

Total Internal Reflection



\* Light travels from denser to rarer medium along OA and gets refracted along AB

\* Increase in angle of incidence the ray 'OA' gets refracted along A, B1.

\* still Increase in angle of Incidence OA2 gets

refracted along  $A_2B_2$  so that  $\angle u = 90^\circ$ . The angle of Incidence is called **critical angle**.

\* still increase in angle of incidence i.e.  $\angle i > c$  (greater than critical angle), now the ray  $OA_3$  gets refracted along  $A_3B_3$ . This is called **Total internal reflection**.

\* Define critical angle?

It is the angle of incidence in the denser medium for which angle of refraction in rarer medium is  $90^\circ$ .

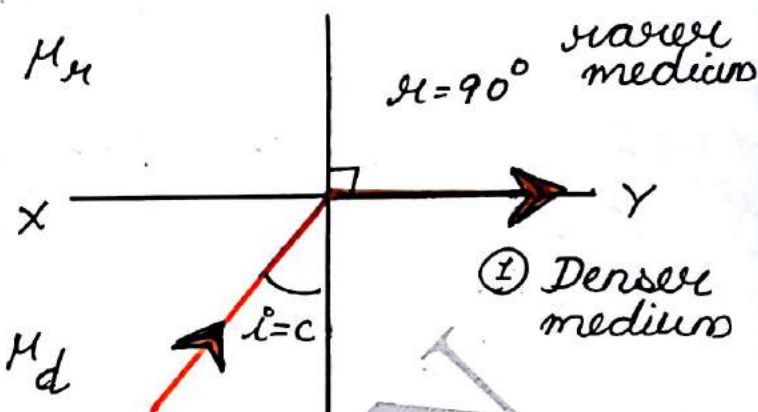
\* Define Total Internal reflection?

When light travels from denser medium to rarer medium and if  $\angle i$  is greater than critical angle, the light ray is reflected back internally without any loss of light.

\* Conditions for Total Internal reflection

- 1, Light ray should travel from denser to rarer medium.
- 2, angle of Incidence should be greater than critical angle.

Derive  $\mu = \frac{1}{\sin c}$



According to Snell's law

$$\frac{\sin i}{\sin r} = \mu_{rd}$$

$$\frac{\sin c}{\sin 90^\circ} = \mu_{rd}$$

$$\frac{\sin c}{1} = \mu_{rd}$$

$$\mu_{dr} = \frac{1}{\sin c}$$

Write the practical applications of total Internal reflection?

- 1, Total internal reflecting prisms
- 2, Mirage
- 3, Optical fibre
- 4, Brilliance of Diamonds.

1, Total Internal reflecting prisms.

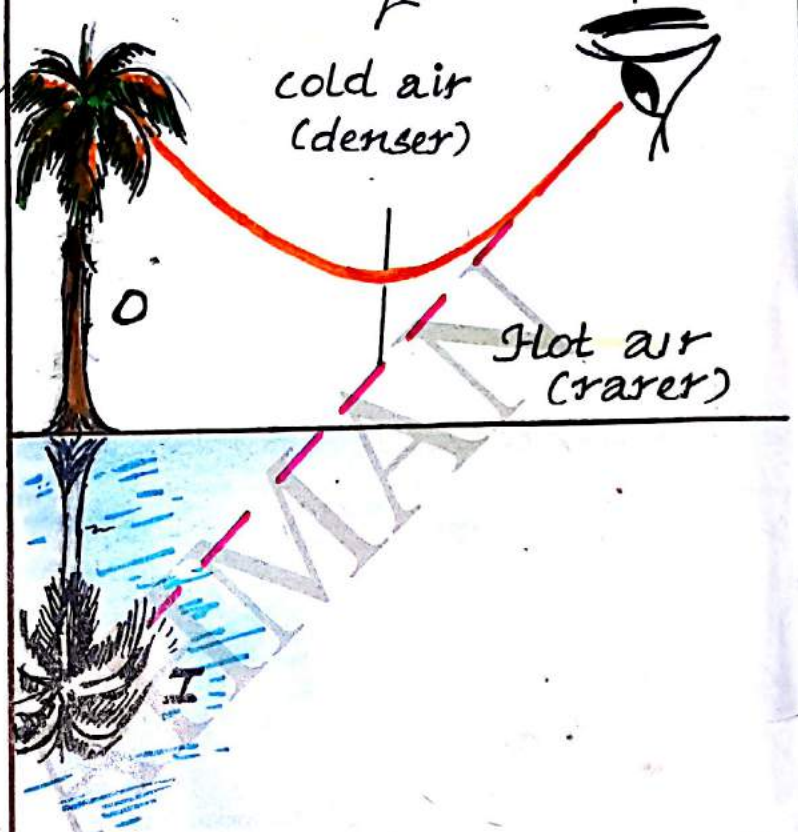
Critical angle for glass is  $42^\circ$ . If a ray is incident at a



glass surface' at an angle greater than  $42^\circ$ , the ray is reflected internally.

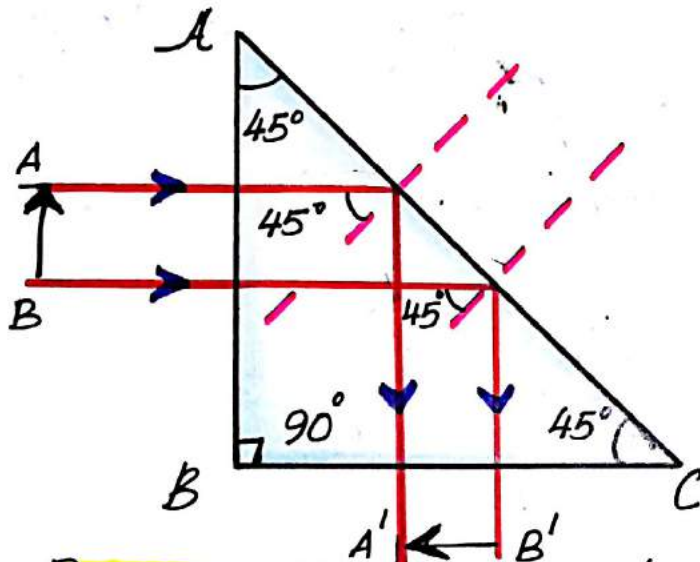
# Mirage

principle :- Total Internal Reflection.

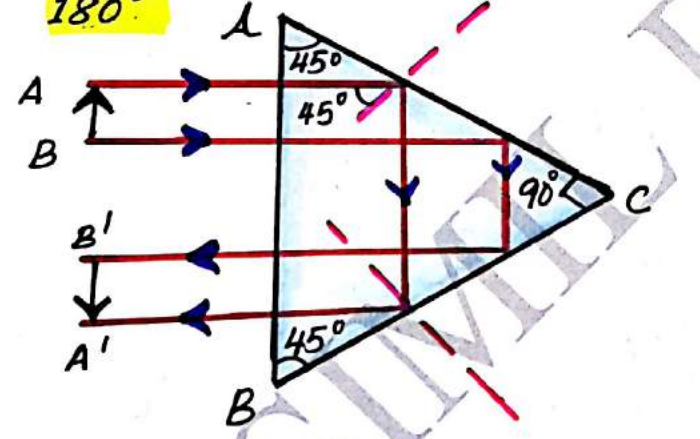


- \* It is an optical illusion.
- \* In hot deserts the object 'O' appears inverted and the observer gets an impression formed by a pond of water.
- \* when light from object 'O' travels from denser medium (cool air well above the ground) to the rarer medium (hot air near ground) it bends away from normal.
- \* when  $\angle i > \angle c$ , it undergoes total internal reflection, and forms an inverted image.

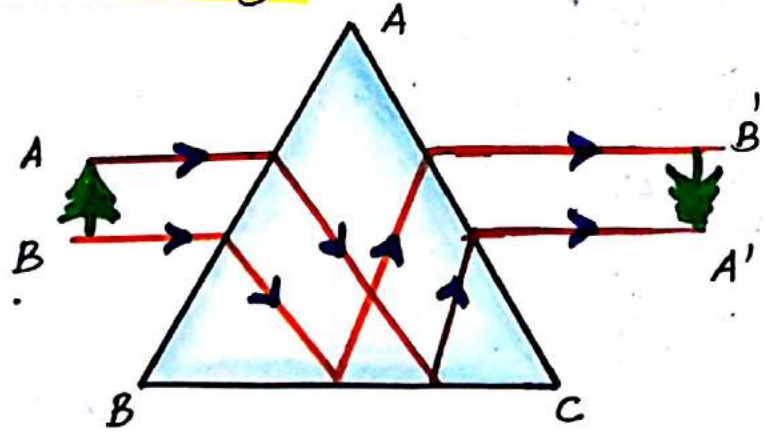
## (a) Ray deviation through $90^\circ$



## (b) Ray deviation through $180^\circ$



## (c) Formation of Inverted image.

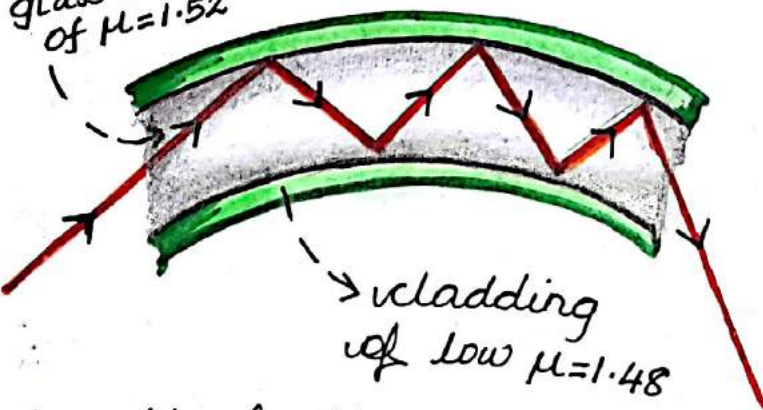


## Optical fibre

\* The principle used in optical fibre is total internal reflection



glass core  
of  $\mu = 1.52$



\* optical fibre consists of a core of high refractive index and a cladding of low refractive index.

\* when light enters optical fibre at suitable angle, it undergoes multiple total internal reflection without any loss of light and energies.

### uses of optical fibre

- 1, communication - Transmission of audio and video signals.
- 2, medical purpose - For visual examination of internal organs like stomach, intestine etc
- 3, In decorative lamps.

## Brilliance of Diamond<sup>5</sup>

\* critical angle for diamond is  $24.4^\circ$

\* A ray entering Diamond undergoes multiple total internal reflection as a result it shines brilliantly



Advantages of total internal reflecting prism over the silvered plane mirror.

- \* silvering is not required
- \* 100% reflection can be achieved
- \* Image will be brighter
- \* Long lasting.

## SPHERICAL REFRACTING SURFACES..

A refracting surface which forms a part of sphere of transparent refracting material.

Relation b/w  $u, v, n_1$  &  $R$

The figure shows the formation of image  $I$  of an object  $O$  on the principal axis of a spherical surface  $XY$  with the centre of curvature  $C$  and radius of curvature  $R$ .

A ray  $ON$  incident from a medium of refractive index  $n_1$  to a medium of refractive index  $n_2$  and it gets refracted along  $NI$ . From the diagram

$$\left. \begin{aligned} i &= \alpha + \gamma \\ r &= \gamma - \beta \\ n &= \gamma - \beta \end{aligned} \right\} \dots (1)$$

$$\left. \begin{aligned} \tan \alpha &= \frac{MN}{MO} \\ \tan \beta &= \frac{MN}{MI} \\ \tan \gamma &= \frac{MN}{NC} \end{aligned} \right\}$$

For small angles

$$\left. \begin{aligned} \tan \alpha &= \alpha = MN/MO \\ \tan \beta &= \beta = MN/MI \\ \tan \gamma &= \gamma = MN/MC \end{aligned} \right\} \dots (2)$$

According to Snell's law

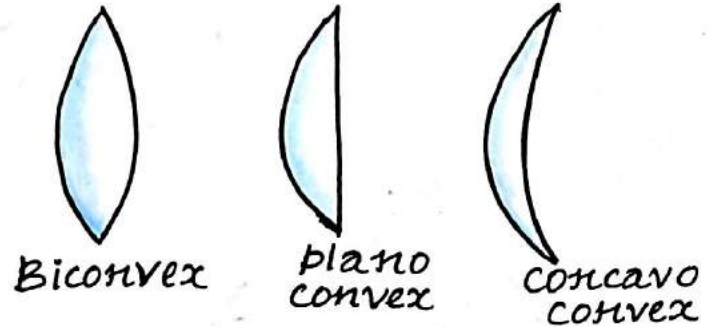
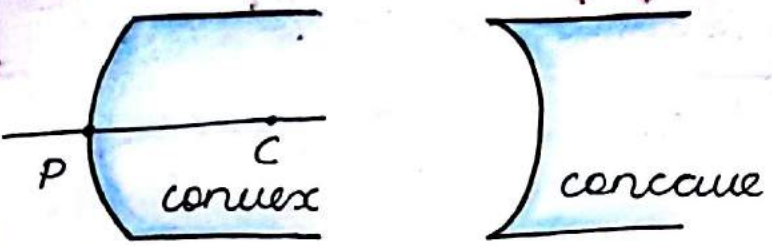
$$\frac{\sin i}{\sin r} = \frac{n_2}{n_1}$$

for small angles  $\sin i = i$   
 $\sin r = r$

$$\therefore \frac{i}{r} = \frac{n_2}{n_1}$$

$$n_1 i = n_2 r \dots (3)$$

$$\text{Div } (3) \Rightarrow n_1 (\alpha + \gamma) = n_2 (\gamma - \beta)$$

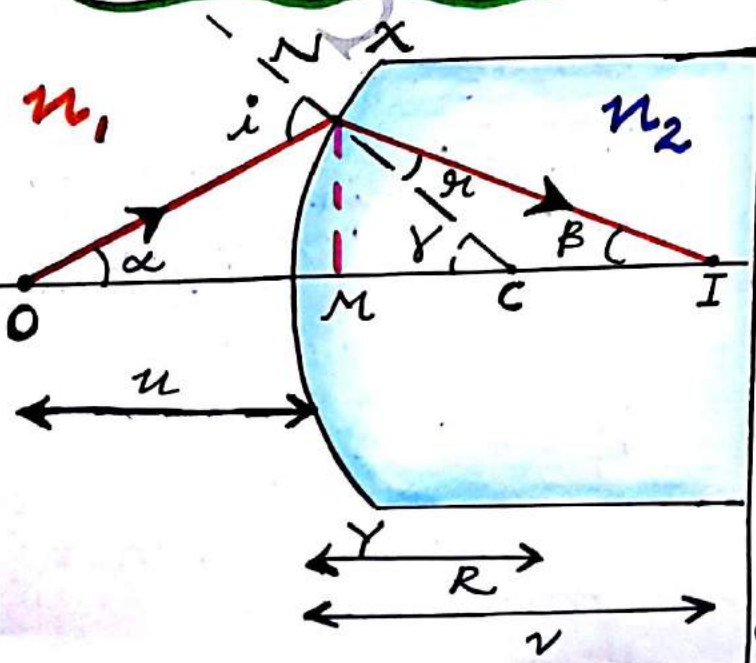


Assumptions to be made in dealing spherical surface.

1. Aperture should be small
2. The object should be a point object
3. Incident ray and refracted ray make small angles with principal axis.

Derive the relation

$$\frac{n_2}{v} - \frac{n_1}{u} = \frac{n_2 - n_1}{R}$$



② in ③

$$n_1 \left( \frac{MN}{MO} + \frac{MN}{MC} \right) = n_2 \left( \frac{MN}{MC} - \frac{MN}{MI} \right)$$

$$n_1 \left( \frac{1}{MO} + \frac{1}{MC} \right) = n_2 \left( \frac{1}{MC} - \frac{1}{MI} \right)$$

$$\{MO = -u, MC = R, MI = v\}$$

$$\frac{n_1}{MO} + \frac{n_1}{MC} = \frac{n_2}{MC} - \frac{n_2}{MI}$$

$$\frac{n_1}{-u} + \frac{n_1}{R} = \frac{n_2}{R} - \frac{n_2}{v}$$

$$\frac{n_2}{v} - \frac{n_1}{u} = \frac{n_2 - n_1}{R}$$

what is meant by a lens?

It is a portion of transparent medium bound by two spherical surfaces.



Biconcave



plano concave



convexo concave

ASSUMPTIONS

1. Aperture of lens should be small.
2. lens should be thin.
3. Incident ray and reflected ray should make small angles with principal axis.

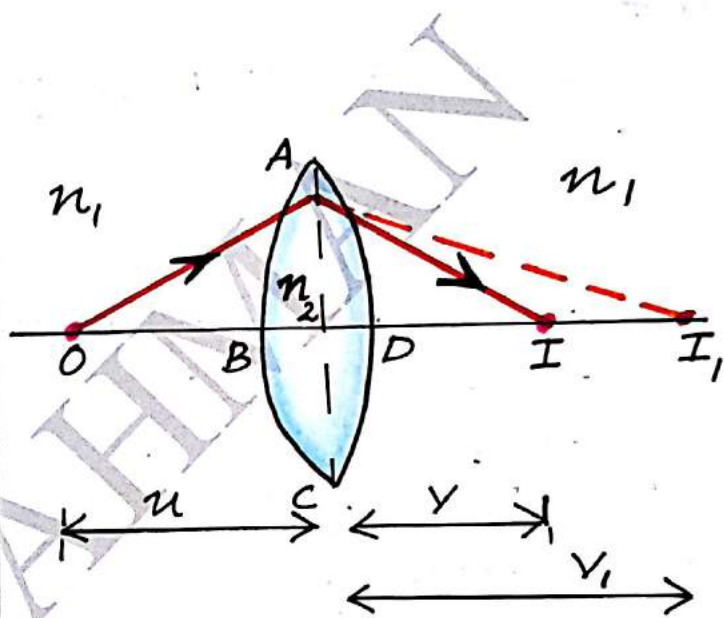
\* How is Intensity related with aperture?

$$\text{Intensity} \propto (\text{Aperture})^2$$

Derive lens-makers formula

$$\frac{1}{f} = (n-1) \left( \frac{1}{R_1} - \frac{1}{R_2} \right)$$

→ Relationship b/w  $f, n, R_1$  and  $R_2$ .



The figure shows the image formation by a double convex lens.

When a light ray incident on the refracting surface ABC, it forms an image I, at a distance  $v_1$

The image I, acts as a virtual object for the formation of image I by the second surface ADC.

at face ABC

$$\frac{n_2}{v_1} - \frac{n_1}{u} = \frac{n_2 - n_1}{R_1} \quad \text{--- (1)}$$

at face ADC

$$\frac{n_1}{v} - \frac{n_2}{v_1} = \frac{n_1 - n_2}{R_2}$$

$$= -\frac{(n_2 - n_1)}{R_2} \dots \textcircled{2}$$

① + ②

$$\frac{n_1}{v} - \frac{n_1}{u} = (n_2 - n_1) \left( \frac{1}{R_1} - \frac{1}{R_2} \right)$$

$\div n_1$

$$\Rightarrow \frac{1}{v} - \frac{1}{u} = \left( \frac{n_2}{n_1} - 1 \right) \left( \frac{1}{R_1} - \frac{1}{R_2} \right)$$

$$\frac{1}{v} - \frac{1}{u} = (n - 1) \left( \frac{1}{R_1} - \frac{1}{R_2} \right) \dots \textcircled{3}$$

If  $u = -\infty$ ,  $v = f$

$$\frac{1}{f} = (n - 1) \left( \frac{1}{R_1} - \frac{1}{R_2} \right) \dots \textcircled{4}$$

This is lens

makes formula.

From ③ and ④

we can say that

$$\frac{1}{v} - \frac{1}{u} = \frac{1}{f}$$

\* Derive the relation  $\frac{1}{f} = \frac{1}{v} - \frac{1}{u}$  for a biconvex lens.

\* Derive laws of distances

\* Derive the relation between object distance ( $u$ ), image distance

( $v$ ) and focal length ( $f$ ) of a lens.

Derive this lens formula?

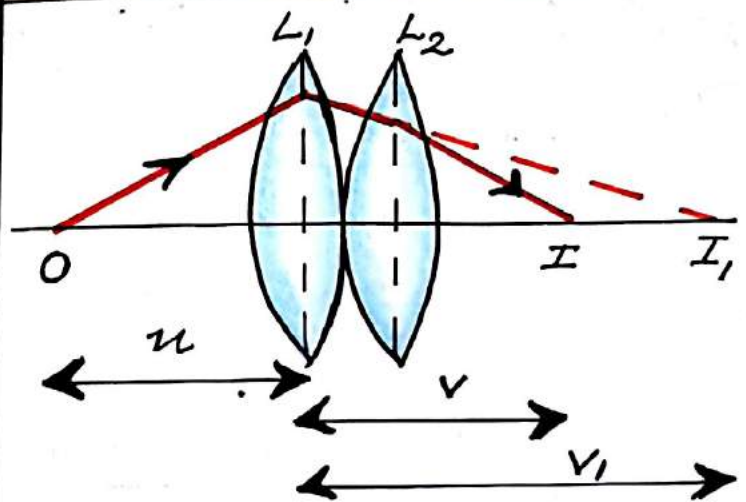
③ = ④

$$\frac{1}{v} - \frac{1}{u} = \frac{1}{f}$$

\* How to estimate rough focal length of a converging lens in lab?

- place lens in front of screen
- adjust position of lens until you get a clear image of a distant object.
- Distance between lens and screen gives rough focal length of lens.

\* Derive expression for the effective focal length of combination of thin lenses?



$f_1$  - focal length of  $L_1$

$f_2$  - focal length of  $L_2$

$L_1, L_2$  are convex lens

A light ray incident on  $L_1$  forms an image  $I_1$  at a distance  $v_1$  from  $C_1$ .

The image  $I_1$  acts as a virtual object for the formation of image  $I$  by  $L_2$ . Distance of  $I$  from  $C_2$  is  $v$ .

For image formed by  $L_1$ ,

$$\frac{1}{v_1} - \frac{1}{u} = \frac{1}{f_1} \dots \dots \textcircled{1}$$

For image formed by  $L_2$

$$\frac{1}{v} - \frac{1}{u_1} = \frac{1}{f_2} \dots \dots \textcircled{2}$$

$$\textcircled{1} + \textcircled{2} \quad \frac{1}{v} - \frac{1}{u} = \frac{1}{f_1} + \frac{1}{f_2}$$

$$\frac{1}{f} = \frac{1}{f_1} + \frac{1}{f_2}$$

\* Focal length of combined lenses.

\* Magnification produced by lens

$$m = \frac{h_i}{h_o} = \frac{v}{u}$$

\* Power of a lens

Ability of a lens to converge or diverge the rays of light.

$$P = \frac{1}{f(m)}$$

$$P = \frac{100}{f(cm)}$$

SI unit of power.  $\neq$  dioptre

\* Magnification 'm' is +ve for virtual image

\* 'm' is -ve for real image.

\* Define one dioptre

$$P = \frac{1}{f}$$

if  $f = 1m$ ,  $P = 1D$

one dioptre is the power of lens when its focal length is 1m.

$$\frac{1}{f} = (n-1) \left( \frac{1}{R_1} - \frac{1}{R_2} \right)$$

$$P = (n-1) \left( \frac{1}{R_1} - \frac{1}{R_2} \right)$$

\* Uses of combined lenses

1. To make final image erect
2. To increase magnification
3. To reduce aberration

\* Expression for power and magnification of combined lenses.

$$\frac{1}{f} = \frac{1}{f_1} + \frac{1}{f_2}$$

$$P = P_1 + P_2$$

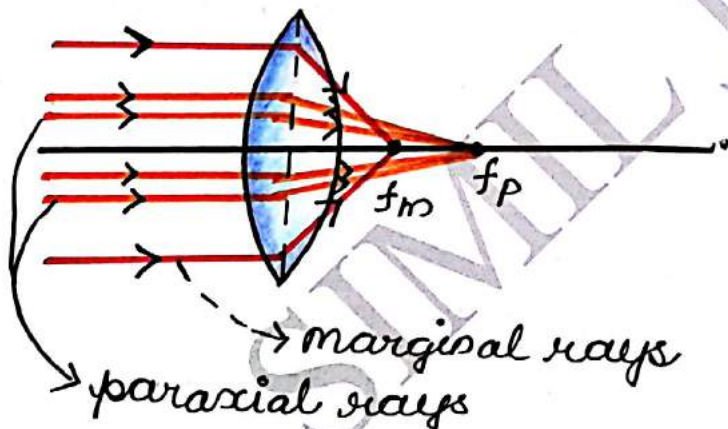
$$m = m_1 \times m_2$$

# Spherical aberration or chromatic aberration

The inability of a lens of large aperture to bring all the rays in a wide beam to focus at a single point is called spherical aberration.

## \* Cause for spherical aberration

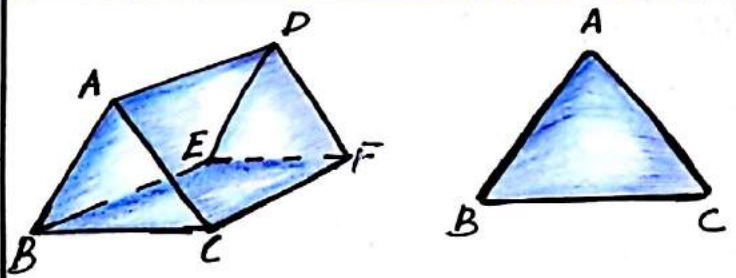
→ Marginal rays focussed closer to optic centre compared to paraxial rays.



## \* How to minimise spherical aberration

- 1. By using stops-slits
- 2. By using combined lenses.

# PRISM

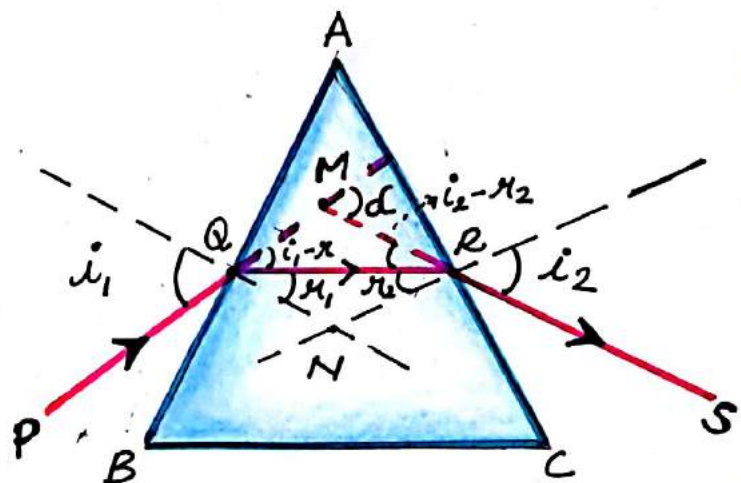


A prism is a portion of a transparent refracting medium bounded by two plane faces inclined to each other at a certain angle.

\* The two plane faces ABED and ACFD are called refracting faces.

\* Angle between them is called angle of prism (A)

$$n = \frac{\sin\left(\frac{A+D}{2}\right)}{\sin\left(\frac{A}{2}\right)}$$



Let ABC be a principal section of a prism with angle of

prism  $A$ .

A ray of light  $PQ$  incident on the face  $AB$  at an angle of incidence  $i_1$  and refracted along  $QR$  at an angle of refraction  $r_1$ .

The ray  $QR$  strikes the face  $AC$  at an angle  $r_2$  and gets refracted along  $RS$  at an angle  $i_2$ .

The angle between the direction of incident ray and emergent ray is called angle of deviation  $d$ .

- \*  $d$  depends upon
- (i) angle of prism ( $A$ )
  - (ii) Material of prism
  - (iii) angle of incidence ( $i$ )

- Figure.
- $ABC$  - section of a prism
  - $PQ$  - Incident ray
  - $QR$  - Refracted ray
  - $RS$  - Emergent ray
  - $QN, RN$  - Normals
  - $A$  - angle to prism
  - $d$  - angle of deviation

In quadrilateral  $AQNR$

$$\angle A + \angle N = 180^\circ \dots \dots \textcircled{1}$$

In  $\Delta QNR$

$$r_1 + r_2 + \angle N = 180^\circ \dots \dots \textcircled{2}$$

compare  $\textcircled{1}$  and  $\textcircled{2}$

$$\angle A = r_1 + r_2 \dots \dots \textcircled{3}$$

In  $\Delta QNR$

sum of interior angles = exterior angle.

$$i_1 - r_1 + i_2 - r_2 = d$$

$$i_1 + i_2 = d + (r_1 + r_2)$$

$$i_1 + i_2 = d + A \dots \dots \textcircled{4}$$

at minimum deviation position ( $D$ )

$$i_1 = i_2 \text{ and } r_1 = r_2 \text{ and } d = D$$

Equation  $\textcircled{3}$  becomes

$$2r = A$$

$$r = \frac{A}{2}$$

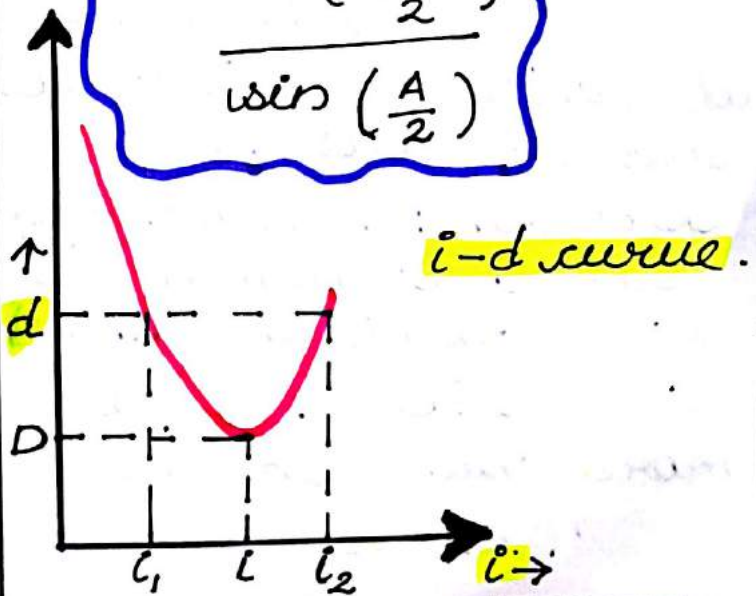
Equation  $\textcircled{4}$  becomes

$$2i = D + A$$

$$i = \frac{A + D}{2}$$

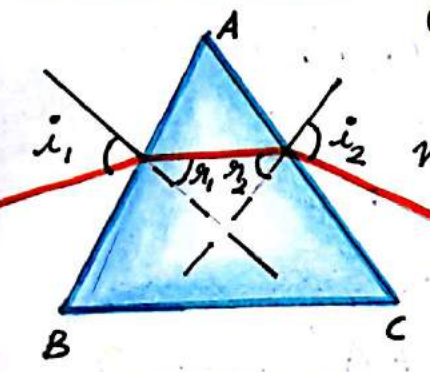
$$n = \frac{\sin i}{\sin r}$$

$$n = \frac{\sin \left( \frac{A + D}{2} \right)}{\sin \left( \frac{A}{2} \right)}$$





angle of deviation for a small angled prism.



at face AB  
 $n = \frac{\sin i_1}{\sin r_1} = \frac{i_1}{r_1}$   
 $i_1 = n r_1$

at face AC  
 $n = \frac{\sin i_2}{\sin r_2} = \frac{i_2}{r_2}$   
 $i_2 = r_2 n$   
 $i_2 = n r_2$

$i_1 + i_2 = A + d$   
 sub  $i_1$  and  $i_2$   
 $n r_1 + n r_2 = A + d$   
 $n (r_1 + r_2) = A + d$   
 $n A = A + d$   
 $d = n A - A$   
 $d = (n - 1) A$

Dispersion of Light

splitting up of white light into constituent colours.

spectrum: The band of colours so formed is called spectrum (VIBGYOR)

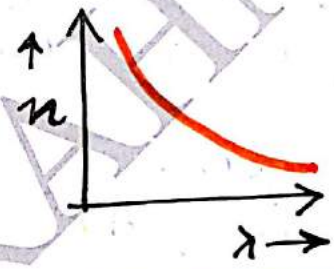
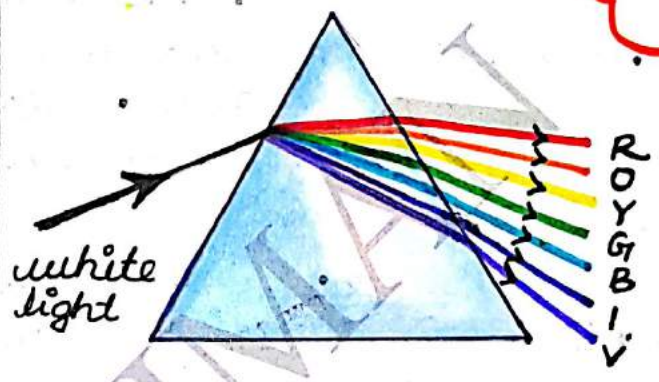
violet ray bends more than red ray  
 $n_v > n_r$

$\lambda_v < \lambda_r$

Velocity of violet < velocity of Red

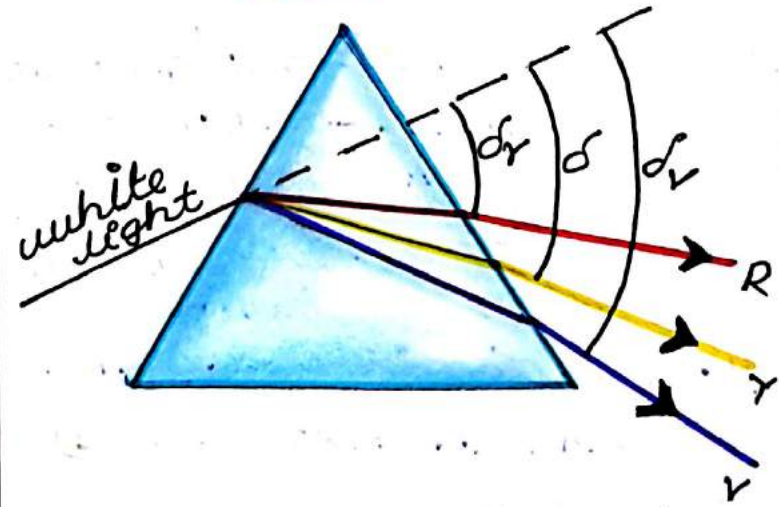
$\delta_{violet} > \delta_{red}$

$n = \frac{c}{v}$   
 $n \propto \frac{1}{v}$



Deviation - Dispersion power

$$W = \frac{n_v - n_r}{n - 1}$$



when white light is passed through a prism, it gets dispersed.

- $d_v$  - deviation of violet colour  
 $d_R$  - deviation of red colour  
 $\delta$  - deviation of mean light (yellow)

### Angular dispersion

$$(d_v - d_R)$$

The difference in angles of deviation of two extreme colours i.e. violet and red colours, is called angular dispersion

$$\delta = (n - 1)A$$

$$d_v = (n_v - 1)A$$

$$d_R = (n_R - 1)A$$

$$d_v - d_R = (n_v - 1)A - (n_R - 1)A$$

$$= n_v A - A - n_R A + A$$

$$d_v - d_R = (n_v - n_R)A$$

Angular dispersion depends on

- (i) angle of prism
- (ii) Nature of material of prism

### Dispersive Power

Ratio between angular dispersion and mean deviation

$$w = \frac{\text{angular dispersion}}{\text{mean deviation}}$$

$$w = \frac{d_v - d_R}{\delta}$$

$$w = \frac{(n_v - n_R)A}{(n - 1)A}$$

$$w = \frac{d_v - d_R}{\delta} = \frac{n_v - n_R}{n - 1}$$

\* Dispersive power depends only on nature of material of prism.

## Scattering of Light

### Rayleigh's scattering law

The intensity of scattered light is inversely proportional to the fourth power of incident light wavelength.

$$I \propto \frac{1}{\lambda^4}$$

\* shorter waves are scattered more than longer waves.

### condition for scattering

size of particle should be in order of  $\lambda$  of light.

### Blue colour of Sky

According to Rayleigh's scattering law  $I \propto \frac{1}{\lambda^4}$

If  $\lambda$  decreases.  $I$  Increases. As blue colour has a shorter wavelength, it gets scattered more by molecules present in atmosphere. So sky appears blue.

Why does sun appear reddish at the time of sunset and sunrise?

- \* At the time of sunset and sunrise, the sun is near the horizon.
- \* Sun light has to travel longer distance.
- \* Most of blue colour scattered away.
- \* Only the least scattered red light reaches the observer.

Why does clouds appear white?

As water droplets size is much greater than  $\lambda$  ( $a \gg \lambda$ ), all wavelengths scattered equally. So cloud appears white.

## Rainbow

The formation of rainbow is due to the dispersion of white light from the sun due to refraction and total internal reflection

from the water droplets suspended in air.

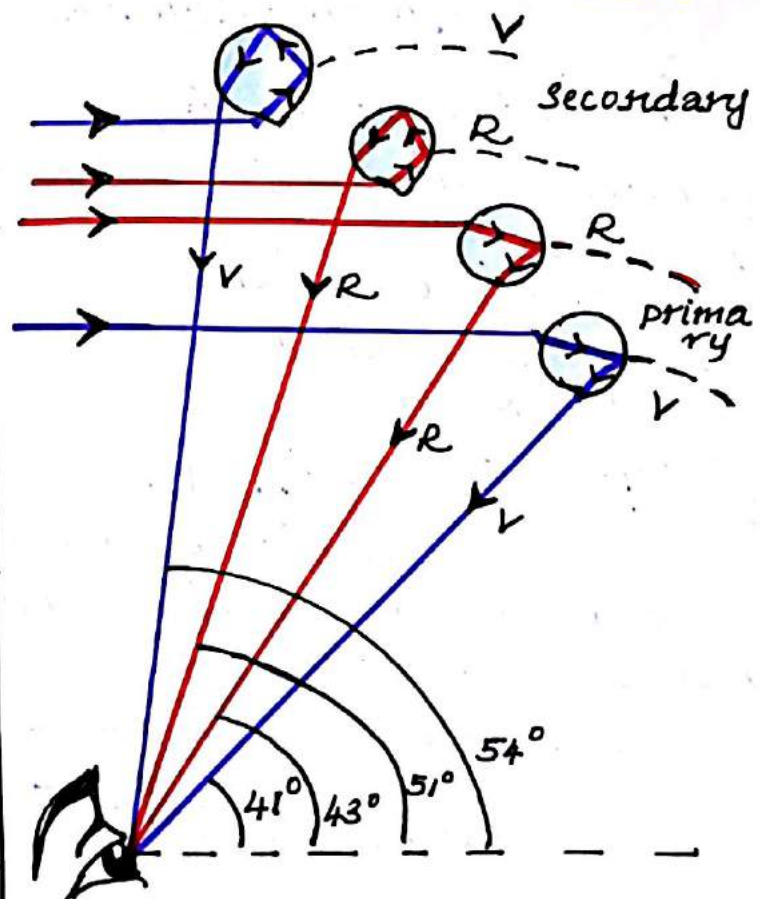
\* Always formed opposite to sun.

Rainbows   
 { Primary Rainbow [brighter one]   
 { Secondary Rainbow [fainter one]

## Primary Rainbow

Formed due to refraction, total internal reflection and dispersion in a droplet.

- \* Outer arc is red and inner arc is violet.
- \* Red emerges at  $42^\circ$  and violet at  $40^\circ$ .



## Secondary Rainbow

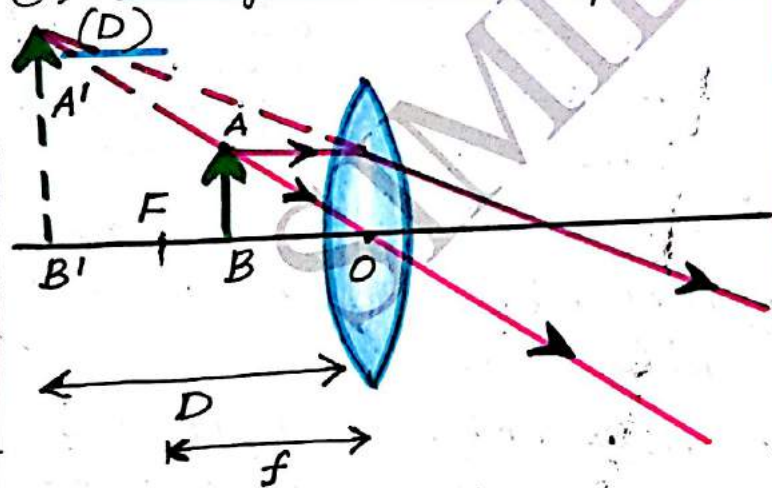
- \* Formed due to refraction, dispersion and two total Internal Reflection within a drop
- \* outer arc is violet & inner arc is red
- \* Red emerges at  $50^\circ$  and violet at  $53^\circ$
- \* It is fainter than primary rainbow.

## Optical Instruments

### Simple microscope

It is a convex lens of small focal length used to get magnified images.

#### (a) Image at near point



when object is placed within the focus, image is adjusted to form at least distance of distinct vision  $D$  ( $D = 25\text{cm}$ )

## Linear magnification (m)

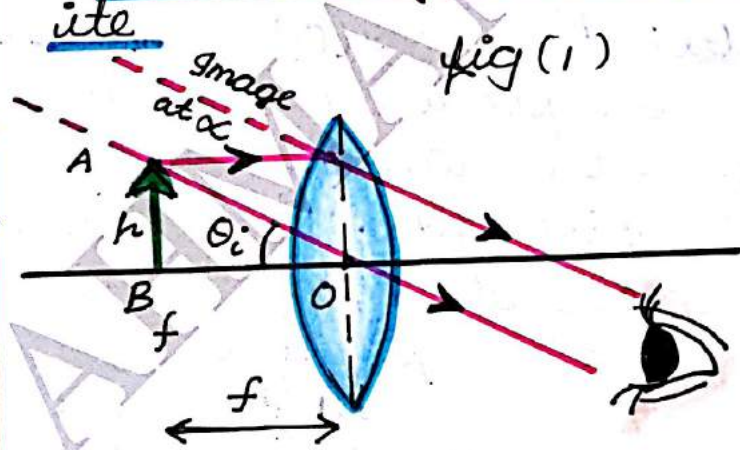
$$m = \frac{v}{u} = v \left( \frac{1}{u} \right) = v \left( \frac{1}{v} - \frac{1}{f} \right)$$

$$m = 1 - \frac{v}{f} \quad \dots \textcircled{1}$$

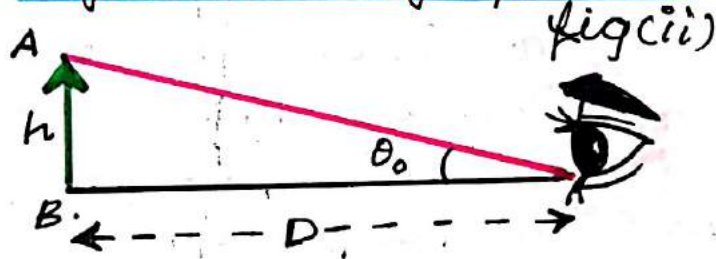
since  $v = -D$

$$\textcircled{1} \Rightarrow m = 1 + \frac{D}{f}$$

#### (b) when image at infinite



### Angular magnification



The maximum angle the object can subtend ( $\theta_o$ ) when object is placed at near point  $D$ . (without convex lens)

$$\tan \theta_o = \frac{h}{D} = \theta_o$$

angle subtended by image (from fig 1)

$$\tan \theta_i = \frac{h}{f} = \theta_i$$

$$m = \frac{\theta_i}{\theta_o} = \frac{h/f}{h/D} = \frac{D}{f}$$

$$m = \frac{D}{f}$$

## Compound Microscope

It has two lenses.

**Objective lens (O)** - very small focal length and short aperture.

**Eye piece lens (E)** - moderate focal length and large aperture.

Place an object **AB** beyond ( $f_o$ ) of objective lens. It forms a real inverted enlarged image **A'B'**. It acts as an object for eye piece, whose position is so adjusted that **A'B'** lies between optic centre  $C_e$  of eye piece and principal focus  $f_e$  of eye piece.

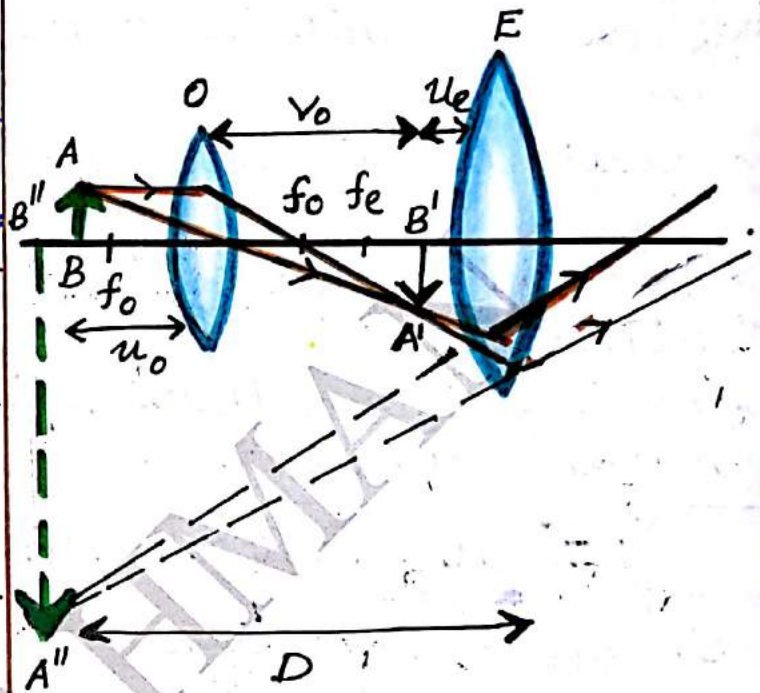
Eye piece forms a virtual, erect and magnified image **A''B''**. It is erect w.r.t **A'B'**.

**AB** - object kept between  $f_o$  and  $2f_o$

**A'B'** - Image formed by O - object for eyepiece

- Real inverted

**A''B''** - Image formed by eye piece at D.



Magnification by objective lens ( $m_o$ )

$$m_o = \frac{A'B'}{AB} = \frac{v_o}{u_o}$$

$$u_o \approx f_o$$

$$v_o \approx L$$

$$\therefore m_o = \frac{L}{f_o}$$

(distance between objective and eye piece = L)

Magnification by eye piece ( $m_e$ )

$$m_e = 1 + \frac{D}{f_e}$$

(proved in simple microscope)

Total magnification ( $m$ )

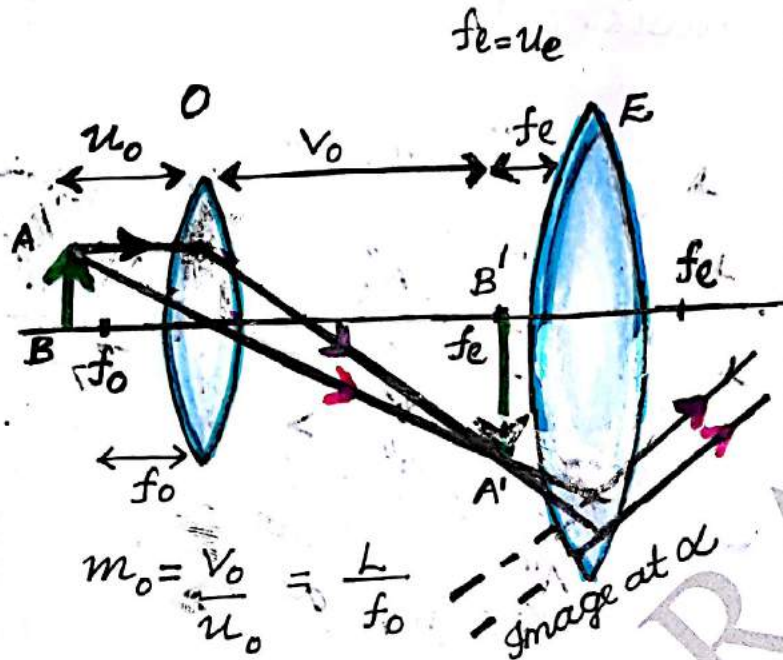
$$m = m_o m_e$$

$$m = \frac{L}{f_o} \left( 1 + \frac{D}{f_e} \right)$$

$$m = \frac{V_o}{u_o} \left( 1 + \frac{D}{f_e} \right)$$

Distance between two lenses  $L = V_o + u_e$

Image formation at Infinity



$$m_o = \frac{v_o}{u_o} = \frac{L}{f_o}$$

$$m_e = \frac{D}{f_e}$$

$$m = m_o m_e$$

$$m = \frac{V_o D}{u_o f_e}$$

$$m = \frac{L D}{f_o f_e}$$

$$L = V_o + f_e$$

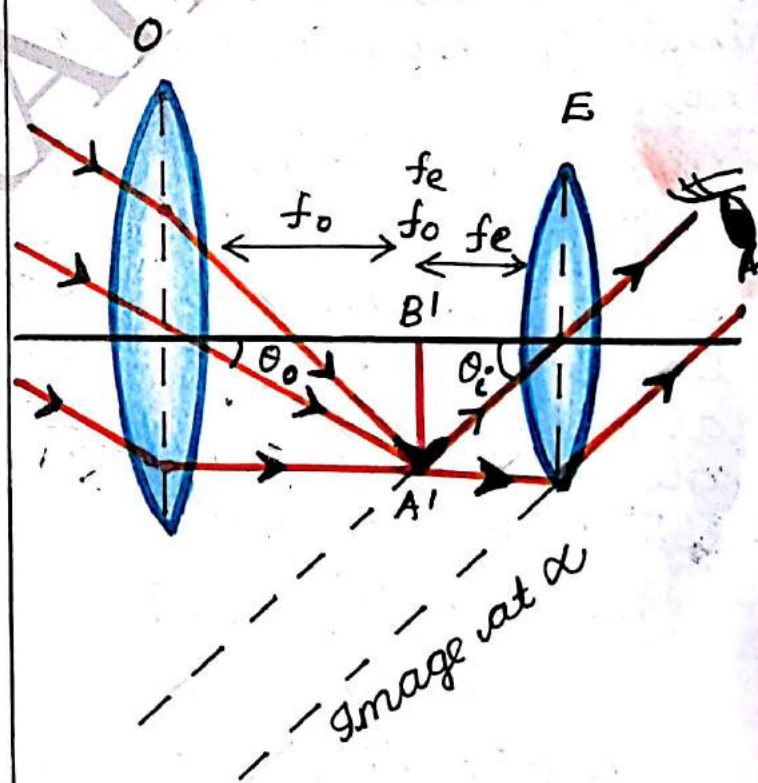
$$L = V_o + u_e$$

to observe very distant objects. (11)

It consists of two lenses, the objective lens: of large focal length and large aperture

eye piece lens: - small focal length and small aperture.

They are fitted at the ends of a tube and the distance between them is adjustable.



Telescope - Astronomical (Refracting type)

Normal adjustment position - Image formation at infinity

It is an optical instrument used

O - objective lens of large focal length. focuses distant object.

E - eye piece of small focal length.

$A'B'$  - Image formed by O.

- object for eye piece
- falls at  $f_o$  as well as at  $f_e$

$\theta_o$  - angle subtended by the object

$\theta_i$  - angle subtended by image.

working:- light from distant object forms a real image  $A'B'$  at the focal plane of the objective lens. This image is adjusted to be at the focal plane of eye piece. Final image is formed at infinity (Normal adjustment)

### Magnification (m)

It is the ratio of the angle subtended by the image at the eye to the angle subtended by the object at the eye.

$$m = \frac{\tan \theta_i}{\tan \theta_o} \dots \textcircled{1}$$

$$\tan \theta_i = A'B' / f_e$$

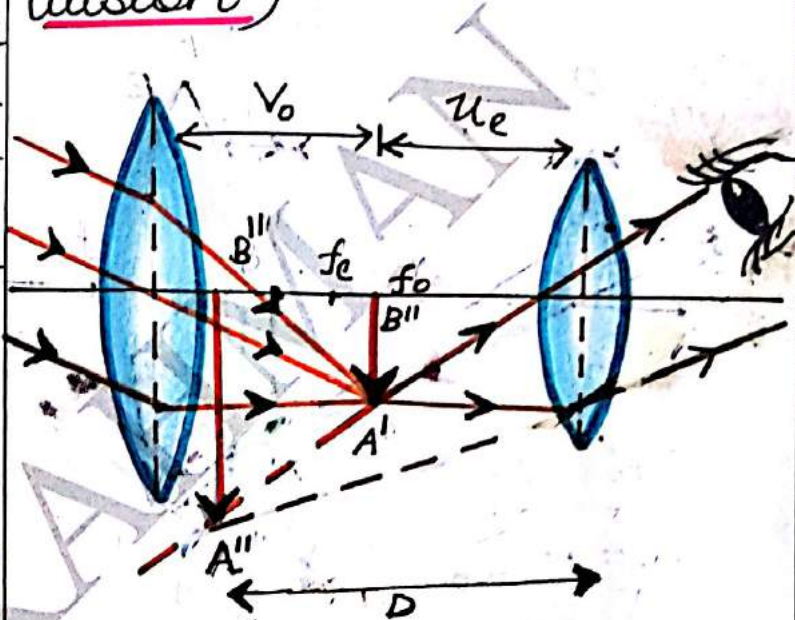
$$\tan \theta_o = A'B' / f_o$$

$$\textcircled{1} \Rightarrow m = \frac{A'B' / f_e}{A'B' / f_o} = \frac{f_o}{f_e}$$

$$m = \frac{f_o}{f_e}$$

$$L = f_o + f_e$$

Image formation at near point (D) - [least distance of distinct vision]



Magnification (final image at D)

$$m = \frac{f_o}{f_e} \left(1 + \frac{f_e}{D}\right)$$

$$L = v_o + u_e$$

### Astronomical Telescope Reflecting Type

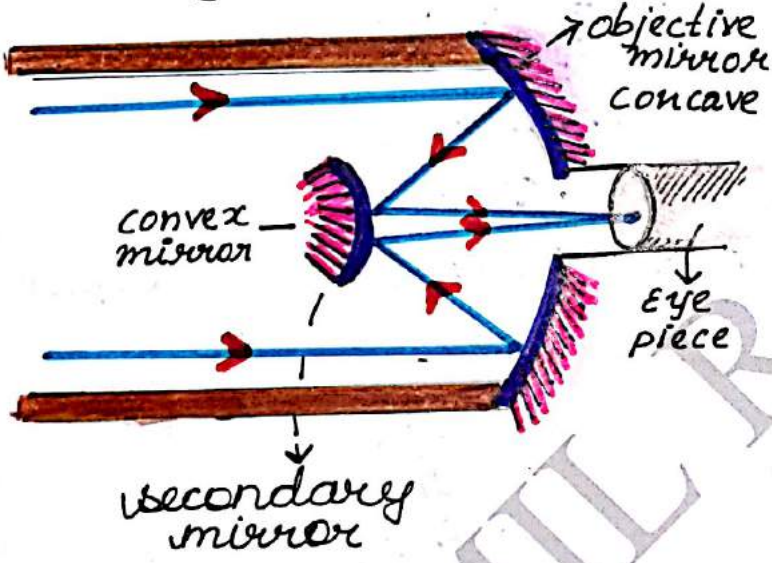
It consists of a concave mirror as objective. parallel rays from distant star are focussed at the principal focus of the concave mirror. The focussed rays fall on a con

convex mirror kept at the principle focus F.

\* Again rays are reflected by convex mirror and focussed to eye piece.

\* Image can be seen through eye piece.

\* Image is inverted, but it is not a matter for astronomical object as they are spherical.



$$m = \frac{f_o}{f_e} = \frac{R/2}{f_e}$$

R - radius of curvature of concave mirror.

Advantages of Reflecting type telescope over refracting type.

- 1, spherical aberration can be minimised.
- 2, large gathering power.
- 3, image is bright
- 4, less weight

Mirror

(a)  $f = R/2$

(b)  $\frac{1}{v} + \frac{1}{u} = \frac{1}{f}$

(c)  $m = \frac{h_i}{h_o} = \frac{-v}{u}$

Refraction - Snell's law

\*  $\frac{\sin i}{\sin r} = n_{21} = \frac{n_2}{n_1} = \text{const}$

\* absolute refractive index

$n = \frac{c}{v}$

\* Refractive Index

$n_{21} = \frac{n_2}{n_1} = \frac{v_1}{v_2} = \frac{\lambda_1}{\lambda_2} = \frac{1}{n_{12}}$

→ Lateral shift

$d = \frac{t \sin(i-r)}{\cos r}$

→ Normal shift

$S = t(1 - \frac{1}{n})$

→ Total Internal reflection

$n = \frac{1}{\sin c}$

→ spherical surface

$\frac{n_2}{v_1} - \frac{n_1}{u} = \frac{n_2 - n_1}{R}$

→ Lens makers formula

$\frac{1}{f} = (n-1) (\frac{1}{R_1} - \frac{1}{R_2})$

$P = (n-1) (\frac{1}{R_1} - \frac{1}{R_2})$

→ Thin lens

\*  $\frac{1}{v} - \frac{1}{u} = \frac{1}{f}$

$m = \frac{h_i}{h_o} = \frac{v}{u}$



→ Lens combination

$$\frac{1}{f} = \frac{1}{f_1} + \frac{1}{f_2}$$

$$P = P_1 + P_2$$

$$m = m_1 \times m_2$$

→ Power

$$P = \frac{1}{f} = \frac{100}{f(\text{cm})}$$

→ Refraction through prism

$$i_1 + i_2 = d + A$$

$$r_1 + r_2 = A$$

@ D  $i_1 = i_2 = i, r_1 = r_2 = r$

$$i = \frac{A + D}{2}, r = A/2$$

$$\therefore n = \frac{\sin(A + D/2)}{\sin(A/2)}$$

→ Deviation for small angled prism

$$d = (n - 1)A$$

→ Dispersive power

$$w = \frac{d_v - d_r}{d} = \frac{n_v - n_r}{n - 1}$$

→ Rayleigh's scattering law

$$I \propto \frac{1}{\lambda^4}$$

→ Simple microscope

(i) Image at D

$$m = 1 + \frac{D}{f}$$

(ii) Image at infinity

$$m = \frac{D}{f}$$

→ Compound microscope

(i) image at D

$$m = \frac{V_o}{u_o} \left(1 + \frac{D}{f_e}\right)$$

$$m = \frac{L}{f_o} \left(1 + \frac{D}{f_e}\right)$$

$$L = V_o + u_e$$

(ii) Image at infinity

$$m = \frac{V_o}{u_o} \frac{D}{f_e}$$

$$m = \frac{L}{f_o} \frac{D}{f_e}$$

$$L = V_o + f_e$$

$$L = V_o + u_e$$

→ Astronomical Telescope

(i) Image at  $\infty$

$$m = \frac{f_o}{f_e}$$

$$L = f_o + f_e$$

(ii) Image at D

$$m = \frac{f_o}{f_e} \left(1 + \frac{f_e}{D}\right)$$

$$L = V_o + u_e$$

→ Telescope - Reflecting Type

$$m = \frac{f_o}{f_e} = \frac{R/2}{f_e}$$